

Absorbing Boundary Conditions for the Finite-Difference Time Evolution of the Wigner Function using the Vlasov Equation

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absorbing boundary conditions. Obviously, the wave package is absorbed at the

[3] T. B. Materdev and C. F. Seyler, "The Quantum Wigner Function in a Magnetic

Field", International Journal of Modern Physics B, 17(25), p4555(2003)

ABSORBING BOUNDARY CONDITIONS SIMULATIONS WITHOUT ABSORBING From the probability density graph we can see it more clearly ABSTRACT BOUNDARY CONDITIONS^[2] $\Psi(x,t)^2$ USING PLANE WAVE DISPERSION^[1] Absorbing boundary conditions for the finite-difference time Given the one-dimensional Schrödinger equation : evolution of the Wigner function using a constant flux condition through $i\hbar \frac{\partial}{\partial x} \Psi(x,t) = \left[-\frac{\hbar^2}{2\pi t^2} \frac{\partial^2}{\partial x} + U(x)\right] \Psi(x,t)$ 100.000 (Th • the 2D region of calculation in phase space are presented. Numerical we first consider plane wave solutions 1000 invision ٠ • results on the time evolution of a Gaussian wave packet in phase $\Psi(x, t) = e^{-i(xt-kt)}$ (8) 150m incation space will be presented and discussed. The conditions minimize The dispersion relation of the solutions is After using absorbing boundary TOTAL INCOME. $hk = \pm [2m^{*}(hw - U)]^{\frac{1}{2}}$ 1st iteration 50th iteration 100th iteration conditions, the reflection is very undesirable reflections at the artificial boundaries of the computation. (9) where plus and minus signs means the right-going and left-going waves, after a linear small (less than 1%) approximation. $\hbar k = \pm \sqrt{2m^2/a} (\hbar w - U)$ (10) Small reflection still exists due to the use of first order finite difference approximation These two relations (9) and (10) intersectuat - U = a BACKGROUND by using $\frac{\partial}{\partial x} \ominus -i\omega$, $\frac{\partial}{\partial x} \ominus -i\omega$, we can rewrite the linear relation as $i\hbar \frac{\partial}{\partial x}\Psi(x,t) = (-i\hbar \sqrt{\frac{a}{2-t}}\frac{\partial}{\partial x} + U)\Psi(x,t)$ 150th iteration dЪ 200th iteration substituting this to the Wigner function, we can get Phase-space density holes are vortex-like nonlinear structures The figures above shows how Wigner function changes in time without applying that have been observed in the magnetosphere. To study the time $\frac{\partial}{\partial x}W(x, v, t) = -\sqrt{\frac{a}{2 - v^2}}\frac{\partial}{\partial y}W(x, v, t)$ (12) absorbing boundary conditions, it reflects at the boundary and quickly diffuses to evolution of these structures we evolve the Wigner function in time or we can write it as the whole space. by solving the Vlasov equation using finite-differences. Wigner $\frac{\partial}{\partial x}W(x,v,t) = -V \frac{\partial}{\partial x}W(x,v,t)$ (13) From Wigner function we can get x probability distributions: function is a quasi-probability density function in phase space. these are the absorbing boundary conditions for the x boundaries, similarly we can get Initial wave package $|\Psi(x,t)|^2 = \int_{0}^{\infty} W(x,v,t) dv$ (2h absorbing boundary conditions for v boundaries. $\frac{\partial}{\partial t}W(x,v,t) = -A \frac{\partial}{\partial t}W(x,v,t)$ W(x,r) EOUATIONS City and an In one-dimensional case, the Wigner Function is ABSORBING BOUNDARY CONDITIONS 1000 incide 100 invation Reflection (Ax=1/100_At=0.002 $W(x, v, t) = \int_{0}^{\infty} (e^{-\frac{i\omega n\Lambda}{2}})\psi^{*}(x - \frac{\Lambda}{2}, t)\psi(x + \frac{\Lambda}{2}, t)d\Lambda$ Reflection (Ax= 1/200_At= 0.004 USING CONSTANT FLUX The graph left shows the temporal COR Incal evolution of the Gaussian wave packet By using smaller Δx and Δt , we can reduce the reflection significantly It satisfies the Vlasov equation without absorbing boundary conditions ➤ We assume the flux f(t) that goes through the boundary (N+1)∆x is the flux that went the wave package reflects and diffuses. through NAx with a time shift T: $f(t)|_{t=1,2,3,1,3,1} = f(t-T)|_{t=2,3,3,1}$ (15) Where $T = \Delta x / V$ (16) CONCLUSIONS In the case $\Delta x \rightarrow 0$, we can assume SIMULATIONS WITH ABSORBING (17) $W((N + 1)\Delta x, v, t) = W(N\Delta x, v, t - T)$ CALCULATIONS > It turns out we get the same absorbing boundary conditions using a we can rewrite it in another form: **BOUNDARY CONDITIONS** dispersion relation for plane waves or using a constant flux assumption. $\frac{W\left((N+1)\Delta x,v,t\right)-W\left((N+1)\Delta x,v,t-T\right)}{W\left((N+1)\Delta x,v,t-T\right)}=\Delta x \frac{W\left(N\Delta x,v,t-T\right)-W\left((N+1)\Delta x,v,t-T\right)}{W\left((N+1)\Delta x,v,t-T\right)}$ > Initial Conditions: We start from a Gauss function which is symmetric both with position and > The remaining small reflection is due to the first-order finite difference which means approximation, as we can reduce that by using higher resolution on x, v or t. momentum $\frac{\partial W(x, v, t)}{\partial t} = -V \frac{\partial W(x, v, t)}{\partial x}$ (19) $W(x, v, 0) = e^{\frac{(x-x')^2}{2\sigma x^2}} e^{\frac{(v-v')^2}{2\sigma x^2}}$ • ٠ where V is the group velocity > Calculations: Similarly we have: 1st iteratio 50th iteration 100th iteration By using finite difference approximation, the Wigner function can be $\frac{\partial W\left(x,v,t\right)}{\partial W\left(x,v,t\right)} = -A \frac{\partial W\left(x,v,t\right)}{\partial W\left(x,v,t\right)}$ (20) evolved step by step in time. REFERENCES at v boundary, where A is the group acceleration. $\frac{\partial W(x,v,t)}{\partial W(x,v,t)} = \frac{W(x,v,t + \Delta t) - W(x,v,t)}{\partial W(x,v,t)}$ $\frac{\partial W(x, v, t)}{\partial W(x, v, t)} = \frac{W(x + \Delta x, v, t) - W(x - \Delta x, v, t)}{\partial W(x - \Delta x, v, t)}$ Compare 13 and 14 to 19 and 20, we get the same absorbing boundary conditions [1] T. Shihata "Absorbing boundary conditions for the finite-difference time-domain $\frac{\partial x}{\partial W(x, v, t)} = \frac{W(x, v + \Delta x, t) - W(x, v - \Delta v, t)}{W(x, v + \Delta x, t) - W(x, v - \Delta v, t)}$ using a dispersion relation for plane waves as using a constant flux assumption. calculation of the one dimensional Schrödinger equation", Phys. Rev. B 43(8). p.6760 (1991) 150th iteration 200th iteration [2] S. Rath, T. Materdey, "Time-Evolution of the Wigner function in Phase-Space Using Finite Differences", APS March Meeting, Baltimore, MD (2006) The figures above shows how the Wigner function changes in time by using the

boundary