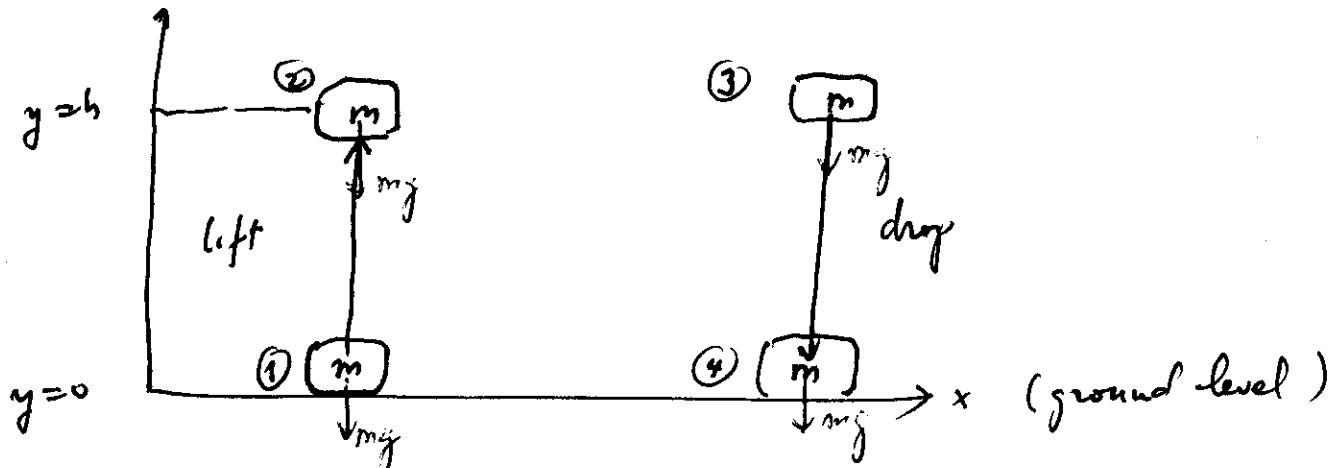


## Ch7: Conservation of Energy

- Work done by a conservative force (e.g.: gravitational force) is conserved
- Work done by a non-conservative force (e.g.: frictional force) is not conserved



Work done by gravity ①  $\rightarrow$  ②

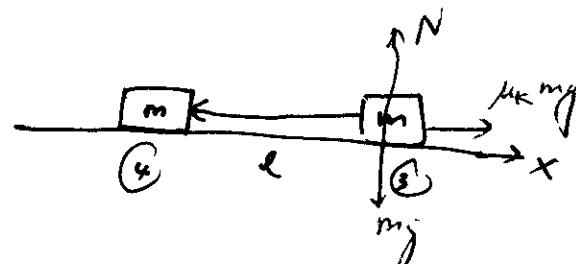
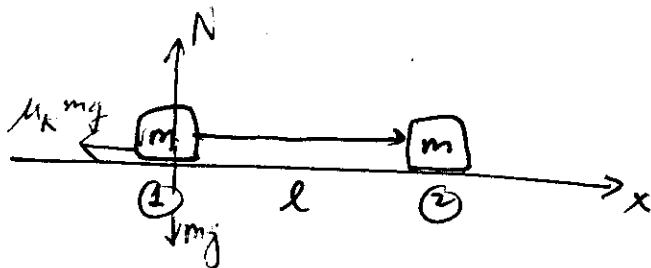
$$\begin{aligned}
 \text{Work}_{(12)} &= \vec{F}_{\text{grav}} \cdot \Delta \vec{y}_{12} \\
 &= -mg\hat{j} \cdot h\hat{j} = -mgh \underbrace{\hat{j} \cdot \hat{j}}_{\perp} \\
 &= -\text{mgh}
 \end{aligned}$$

Work done by gravity ③  $\rightarrow$  ④

$$\begin{aligned}
 \text{Work}_{(34)} &= \vec{F}_{\text{grav}} \cdot \Delta \vec{y}_{34} \\
 &= -mg\hat{j} \cdot (-h\hat{j}) \\
 &= mgh \underbrace{\hat{j} \cdot \hat{j}}_{\perp} = +mgh
 \end{aligned}$$

$\rightarrow$  Total work done by gravity ①  $\rightarrow$  ②  $\rightarrow$  ③  $\rightarrow$  ④ is 0!

- Work done by gravity or gravitational potential energy is conserved (gravity is a conservative force)



Work done by friction:

$$\begin{aligned} \text{Work}_{12} &= \vec{F}_f \cdot \Delta \vec{x}_{12} \\ &= -\mu_k mg \hat{i} \cdot \hat{l} \hat{i} \\ &= -\mu_k mgl \underbrace{\hat{i} \cdot \hat{i}}_{\perp} \\ &= -\mu_k mgl \end{aligned}$$

$$\begin{aligned} \text{Work}_{34} &= \vec{F}_f \cdot \Delta \vec{x}_{34} \\ &= +\mu_k mg \hat{i} \cdot (-\hat{l} \hat{i}) \\ &= -\mu_k mgl \underbrace{\hat{i} \cdot \hat{i}}_{\perp} \\ &= -\mu_k mgl \end{aligned}$$

Total work done by friction:  $\textcircled{1} \rightarrow \textcircled{2} \rightarrow \textcircled{3} \rightarrow \textcircled{4}$  i.e.  $-2\mu_k mgl \neq 0$

- Work done by friction is not conserved. Friction is not a conservative force.

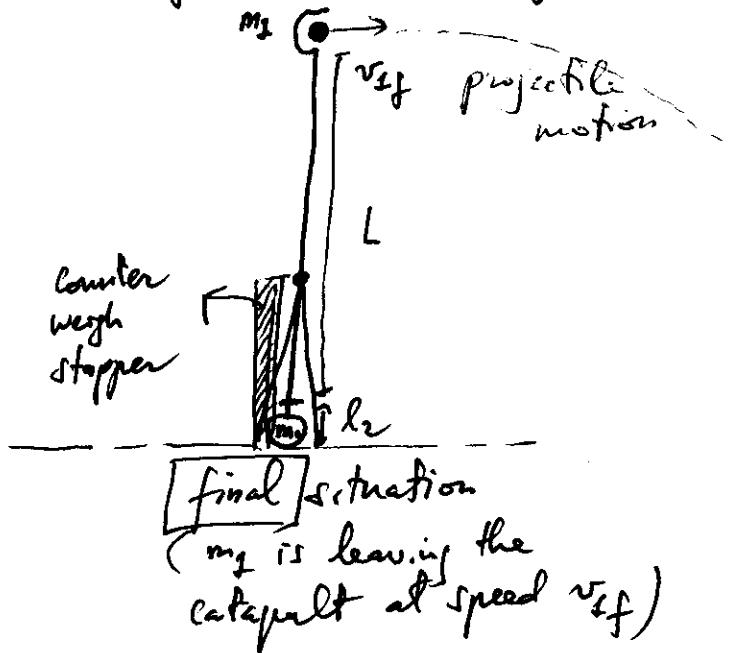
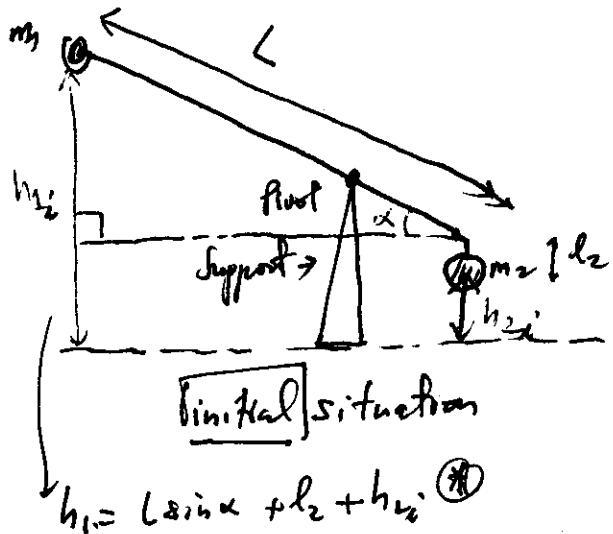
### Conservation of Mechanical Energy:

$\downarrow$                                    $\downarrow$   
 Sum of Kinetic Energy & Grav. Potential Energy  
 is conserved.

$$\underbrace{\frac{1}{2}mv_i^2 + mgh_i}_{\text{initial mechanical energy}} = \underbrace{\frac{1}{2}mv_f^2 + mgh_f}_{\text{final mechanical energy}}$$

PP3 : 7.1

Catapult or Trebuchet: uses a counter weight & launches an object forward (at a higher or lower speed depending on the initial height of the counter weight)  $\rightarrow$  Conservation of Mechanical Energy.



- If we could calculate  $v_{if}$   $\rightarrow$  can find out where  $m_1$  will land based on initial parameters such as  $h_2$ .
  - Limitations: friction at pivot, weight of arm, air resistance.
- With initial & final situations defined  $\rightarrow$  let's apply conservation of mechanical energy:

$$\underbrace{\frac{1}{2}m_1v_{if}^2 + m_2gh_{2i} + \frac{1}{2}m_2v_{if}^2}_{\text{only grav. potential energy at initial}} + m_2gh_{2i} = \frac{1}{2}m_1v_{if}^2 + \frac{1}{2}m_2v_{if}^2 + m_1g(L + l_2)$$

due to stopper

$$\begin{aligned}
 m_1g h_{2i} + m_2 h_{2i} &= \frac{1}{2}m_1v_{if}^2 + m_1g(L + l_2) \\
 m_2 h_{2i} + m_2 h_{2i} &= \frac{1}{2}m_1v_{if}^2 + m_1g(L + l_2 - L \sin \alpha - l_2) \\
 (m_1 + m_2)g h_{2i} &= \frac{1}{2}m_1v_{if}^2 + m_1g L(1 - \sin \alpha)
 \end{aligned}$$

→ This allows us to calculate  $v_{if}!$  from  $m_2, m_c, L, \alpha$

→ projectile motion for  $m_1$  :

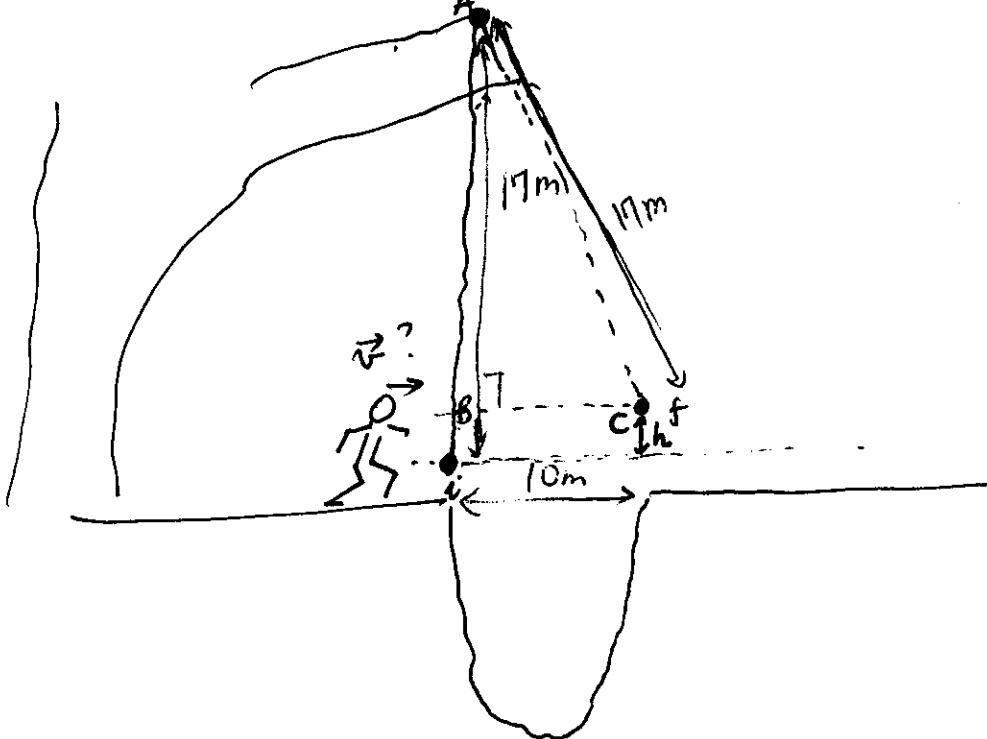
$$\left\{ \begin{array}{l} \text{vertical: constant acceleration} \\ L + h_2 = \frac{1}{2} g t^2 \rightarrow t = \sqrt{\frac{2(L+h_2)}{g}} \end{array} \right.$$

$$\text{horizontal: } x_{\text{range}} = v_{if} \cdot t$$

$$x_{\text{range}} = \sqrt{\frac{2((m_1+m_c)gh_{2i} - m_2gL(1-\sin\alpha))}{m_1}} \cdot \sqrt{\frac{2(L+h_2)}{g}}$$

7.64

80



Tarzan's minimum speed: just enough to land at the other side with no final kinetic energy. Using the initial height (where he grabs on the vine above ground level) as the reference level (zero height), he lands at the other side at height  $h$  (see sketch): this gives a final grav. potential energy into which the initial kinetic energy has converted.

Conservation of ME:

$$\frac{1}{2}mv_i^2 + mgh_i = \frac{1}{2}mv_f^2 + mgh_f$$

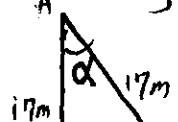
$0$  at minimum

$$\frac{1}{2}mv_{i\min}^2 = mg(h_f - h_i) = mgh$$

$$v_{i\min} = \sqrt{2gh}$$

Calculate  $h$  using geometry (the sketch is essential!)

Alternative #1:  $AB^2 + 10^2 = 17^2 \rightarrow AB = \sqrt{17^2 - 100} = 13.75 \text{ m}$



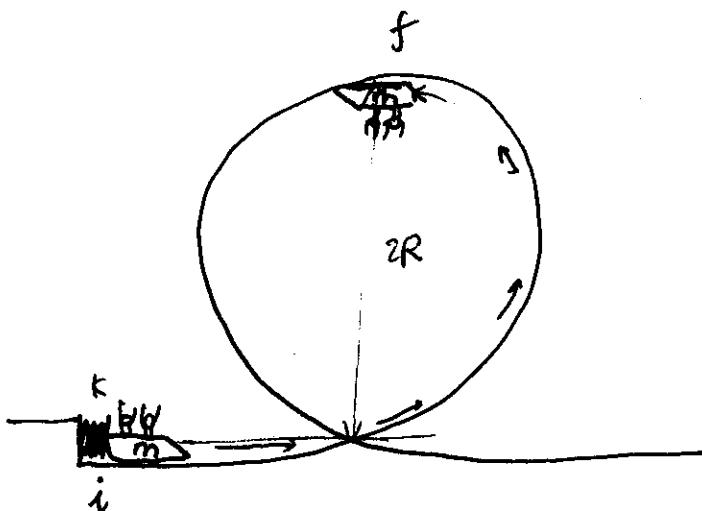
$$\rightarrow h = 17 - AB = 3.25 \text{ m}$$

Alternative #2:  $\sin \alpha = \frac{10}{17} \rightarrow \alpha = \sin^{-1} \frac{10}{17} = 36^\circ$

$$AB = 17 \cos \alpha \rightarrow h = 17 - 17 \cos \alpha / h = 17(1 - \cos \alpha) = 17(1 - \cos(\sin^{-1} \frac{10}{17})) \approx 3.25 \text{ m}$$

$$v_{i\min} = \sqrt{2 \cdot 9.81 \cdot 3.75} = 7.98 \frac{\text{m}}{\text{s}}$$

7.55



$$m = 840 \text{ kg}$$

$$k = 31 \frac{\text{kN}}{\text{m}}$$

$$R = 6.2 \text{ m}$$

Min compression of spring  
for car to make top of loop?

Conservation of mechanical energy:

$$\frac{1}{2}mv_i^2 + mgh_i = \frac{1}{2}mv_f^2 + mgh_f$$

Important:  $v_f \neq 0$  (otherwise it will not make it!)

At f: if we look at min initial speed  $v_{i\min}$ , the only acceleration to change direction of motion in a circular motion is g. In other words: the only agent to change direction of motion at f is mg:

$$mg \frac{v_f^2}{R} = mg g \rightarrow v_f^2 = gR$$

$$\rightarrow \frac{1}{2}mv_{i\min}^2 = \underbrace{\frac{1}{2}mv_f^2}_{\frac{1}{2}mgR} + \underbrace{\frac{mg(h_f - h_i)}{2R}}_{mg^2R} = mg \frac{5}{2}R$$

$$v_{i\min}^2 = 5gR$$

What is  $x_{i\min}^2$  (min compression of spring =)

When a spring is compressed a distance x, elastic potential energy is stored in the amount of:  $\frac{1}{2}kx^2$ . This will go into the initial

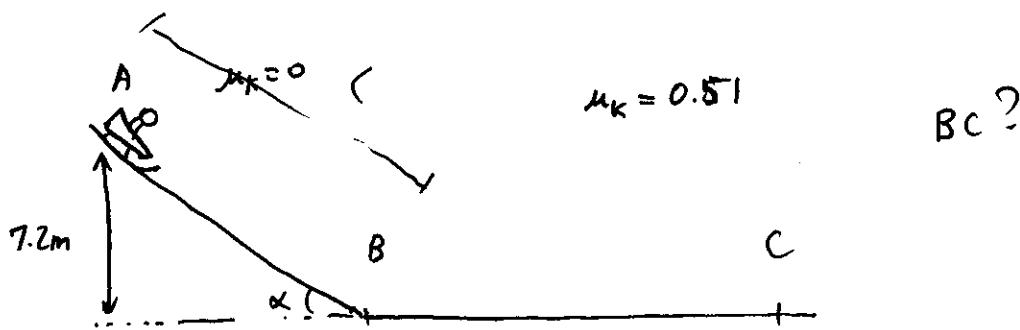
KE of the roller-coaster car:

$$\frac{1}{2} k x_{\min}^2 = \frac{1}{2} m v_{\min}^2 \rightarrow x_{\min} = \sqrt{\frac{m v_{\min}^2}{k}}$$

$$= \sqrt{\frac{m g R}{k}}$$

$$= \sqrt{\frac{840 \cdot 5 \cdot 9.81 \cdot 6.2}{31000}} = 2.87 \text{ m}$$

7.61



AB: constant acceleration

BC: constant deceleration

{ When we solved a similar problem using kinematic equations:

$$v_A = 0 \rightarrow a_{AB} \rightarrow v_B \rightarrow a_{BC}, v_C = 0 \rightarrow BC$$

Note:  $\begin{cases} \alpha \\ L \\ m \end{cases}$  were also given!

Conservation of mechanical energy: only applies to AB  
(since friction b/w B & C is not conservative!). But can find  $v_B$ ,  
then use kinematic equation to find BC.

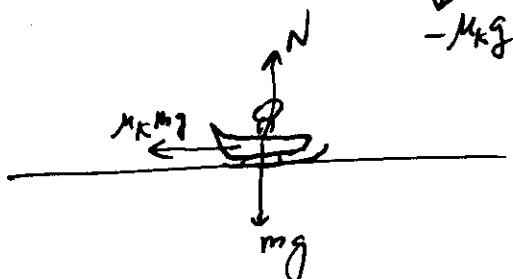
$$\left\{ \begin{array}{l} \text{initial} = @A \\ \text{final} = @B \end{array} \right. \quad \frac{1}{2} m v_A^2 + mgh_A = \frac{1}{2} m v_B^2 + mgh_B$$

$$mgh_A - mgh_B = \frac{1}{2} m v_B^2$$

$$v_B = \sqrt{2g(h_A - h_B)} = \sqrt{2 \cdot 9.81 \cdot \frac{7.2}{2.87}} = 11.9 \text{ m/s}$$

BC: Kinematic equation #3:

$$\frac{v_c^2 - v_B^2}{x_{BC}} = 2 \cdot a_{BC} \rightarrow \frac{-v_B^2}{x_{BC}} = -2 \mu_k g$$



$$x_{BC} = \frac{v_B^2}{2 \mu_k \cdot g}$$

$$= \frac{11.9^2}{2 \cdot 0.51 \cdot 9.81}$$

$$= 14.2 \text{ m.}$$

## Ch 8 : Gravitation

### Universal Law of Gravitation

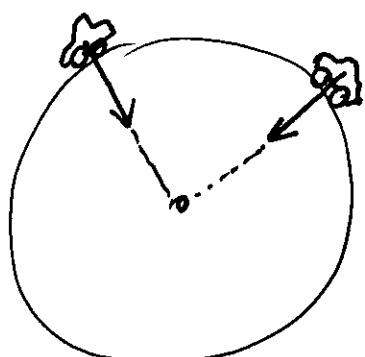
↳ here on Earth, moon, other planets, other galaxies, to the universe

$$F = G \frac{m_1 \cdot m_2}{r^2}$$

Force of grav.  
 attraction b/w  
 $m_1$  &  $m_2$       Universal  
 Grav. Constant:  
 $6.67 \times 10^{-11} \frac{\text{Nm}^2}{\text{kg}^2}$

separation b/w  
 $m_1$  &  $m_2$ !  
 (larger separation  $\rightarrow$   
 less grav. attraction)

Force is a vector; the above is its magnitude. Direction is center to center toward the more massive object.



At our scale: these attractions are vertical & downward  
 (for small distances, the curvature of the Earth is not noticeable)

All along we have been using a force of gravitational attraction of  $mg$ : what is it compared to the Univ. Law of Grav.?

$$(g = 9.81 \frac{\text{m}}{\text{s}^2})$$

$$M_{\text{Earth}} = 5.97 \times 10^{24} \text{ kg} \rightarrow m_1$$

$$R_E = 6.37 \times 10^6 \text{ m} \rightarrow r$$

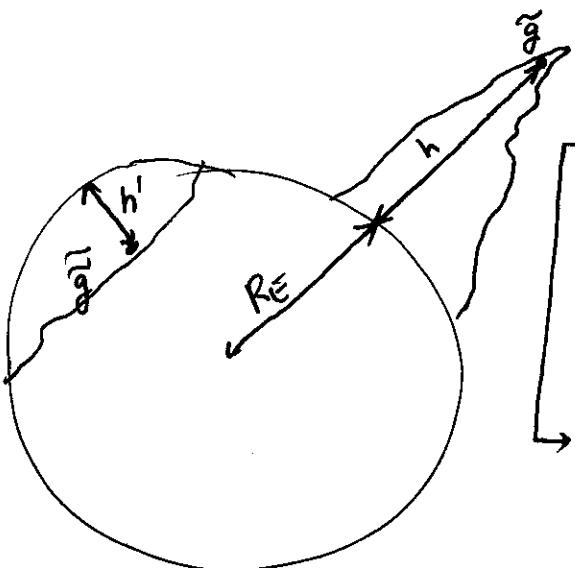
For an object of mass  $m \rightarrow m_2$  on the surface of the Earth: Univ. Law of Grav. says:

$$F = G \frac{M_E \cdot m}{R_E^2} = 6.67 \times 10^{-11} \cdot \underbrace{\frac{5.97 \times 10^{24}}{(6.37 \times 10^6)^2}}_{g=9.81 \frac{\text{m}}{\text{s}^2}}, m$$

$\uparrow$   
mass of object

$$F = m \cdot g$$

$$\text{where } g = G \frac{M_E}{R_E^2}$$



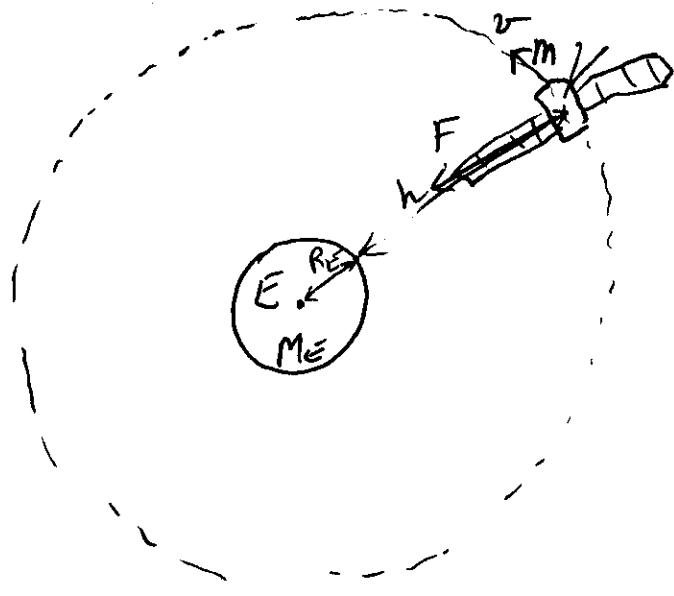
if  $h'$  is not negligible compared to  $R_E \rightarrow \tilde{g} = G \cdot \frac{M_E}{(R_E + h')^2} < g$

at bottom of ocean:  $\tilde{g} = G \frac{M_E}{(R_E - h')^2} > g$

## Orbital Motion:

Circular orbital motion: satellite in UCM (constant speed  $v$ )

- No energy cost to run it in orbital motion. Energy is only needed for operation (signal communication with the planet)
- Energy: from grav. attraction of the Earth:



Force of grav. attraction on satellite is  $F$ , which is essential to keep the satellite in UCM: Why?  
Because it is the agent that changes direction of motion to keep it in orbit and

not to go off in a straight forward direction:

$$F = G \frac{M_E \cdot m}{(R_E + h)^2} = m \cdot \frac{v^2}{R_E + h}$$

$$\rightarrow \text{Orbital speed} \rightarrow v = \sqrt{\frac{G \cdot M_E}{R_E + h}} \quad \left\{ \begin{array}{l} \rightarrow \text{Univ. Law of Grav.} \\ \rightarrow \text{2nd Newton's Law for satellite under UCM (w/ radial acceleration)} \end{array} \right.$$

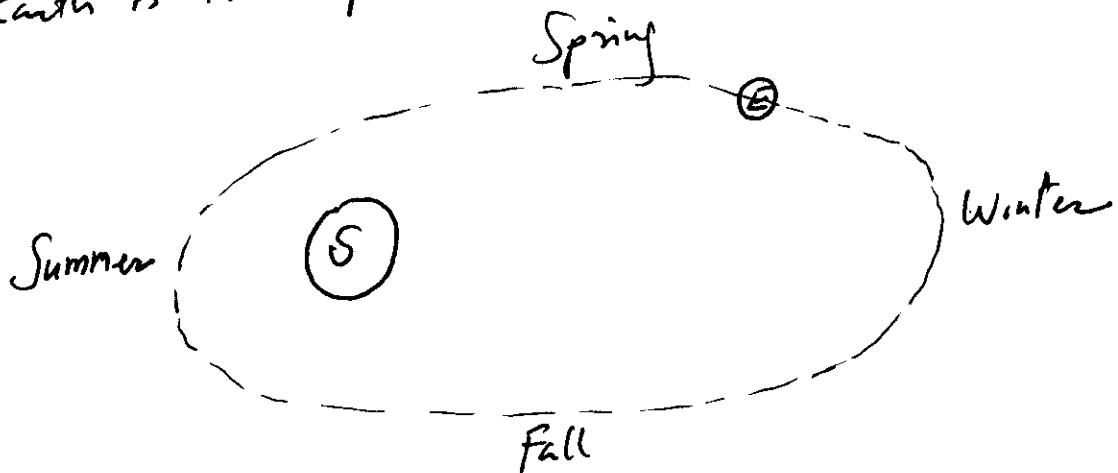
$$\rightarrow \text{Orbital period: time to complete one full orbit:}$$

$$\boxed{T = \frac{2\pi(R_E + h)}{\sqrt{\frac{G \cdot M_E}{R_E + h}}}} = \frac{2\pi}{\sqrt{G \cdot M_E}} (R_E + h)^{3/2} \rightarrow T^2 = \frac{4\pi^2}{GM_E} (R_E + h)^3 \approx R$$

If we extend this relationship b/w period & radius ( $T^2 \propto R^3$ )  
to elliptical orbits  $\rightarrow$

Kepler's 3rd Law: "the period squared is proportional to  
the semimajor axis cubed"

↳ Earth is in elliptical orbit around the Sun:



Cell phone satellite @  $h = 250$  km

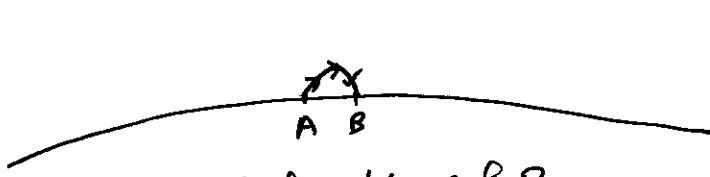
Orbital period for this satellite:

$$T = \frac{2\pi}{\sqrt{GM_E}} (R_E + h)^{3/2} = \frac{2\pi}{\sqrt{6.67 \times 10^{-11} \cdot 5.97 \times 10^{24}}} [(6370 + 250) \cdot 10^3]^{3/2}$$

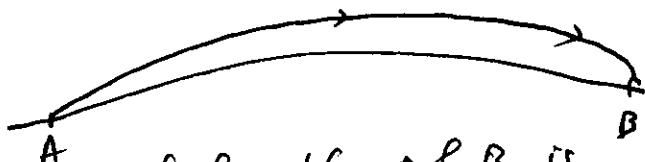
$$= 5400.5 = 1.5h$$

Projectile Motion (in part: vertical component was due to gravitational attraction)

↳ so far its trajectory was a parabola (provided the ground is "flat": balls, bullets, short-range missile)

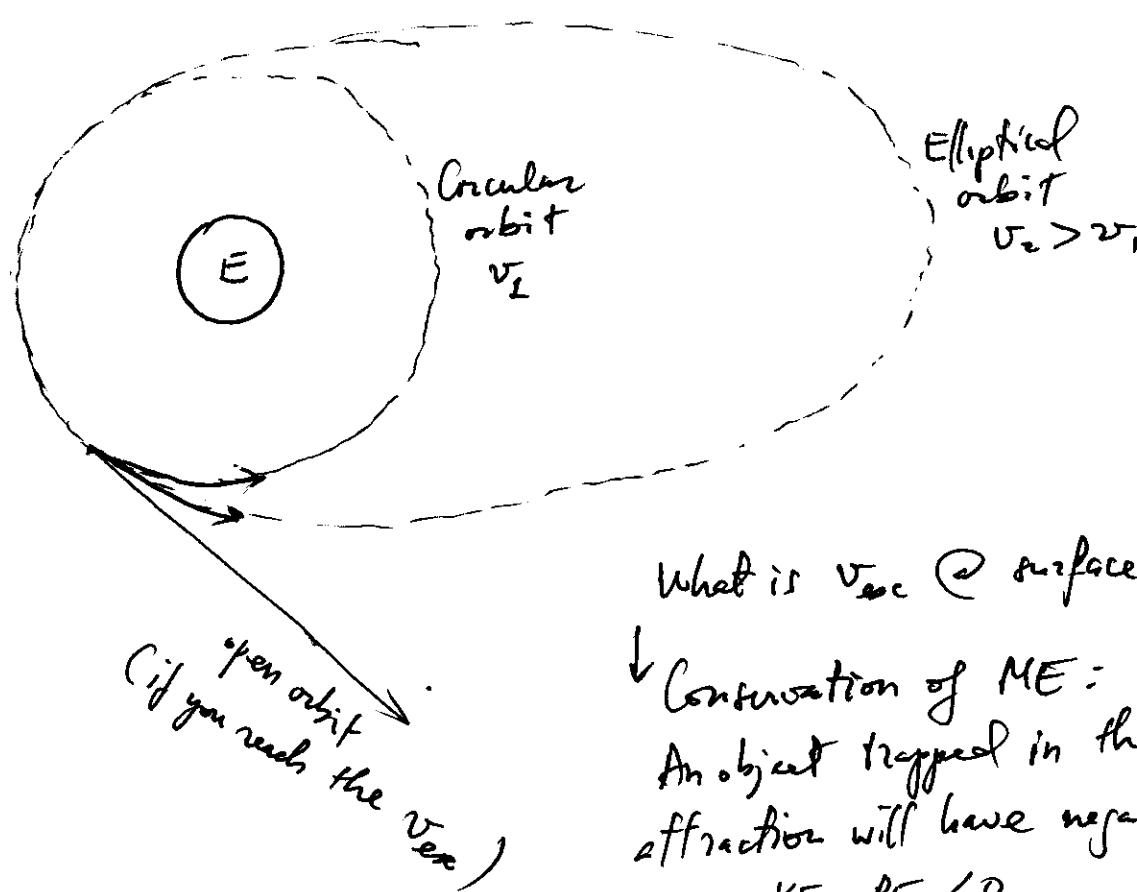


Surface b/w A & B  
is flat to a very  
good approximation  
→ Trajectory is a parabola



Surface b/w A & B is  
no longer flat  
→ Trajectory is part of an  
elliptical orbit  
→ long-range missile, etc.

Escape Speed: if you don't have at least the escape speed  $\rightarrow$  you are trapped in the gravitational attraction of the planet: stuck on surface or in an orbit around the planet. If  $v > v_{esc}$   $\rightarrow$  object will follow an "open orbit" away from the gravitational attraction  $\rightarrow$  can travel to the outer space: interplanetary travel



What is  $v_{esc}$  @ surface of Earth?

↓ Conservation of ME:  
An object trapped in the grav. attraction will have negative ME

$KE + PE < 0$   
At total zero ME  $\rightarrow$  will start open orbit:  
 $v_{esc} \rightarrow ME = 0$

$$\frac{1}{2}mv_{esc}^2 - \frac{GM_E m}{r} = 0 \rightarrow v_{esc} = \sqrt{\frac{2G \cdot M_E}{r}}$$

$$\text{at surface } r = R_E \rightarrow v_{esc} = \sqrt{\frac{2 \cdot 6.67 \times 10^{-11} \cdot 5.97 \times 10^{24}}{6.37 \times 10^6}} \\ = 11.2 \frac{\text{km}}{\text{s}} = 40320 \frac{\text{km}}{\text{h}}$$

Why the grav. potential energy is  $-\frac{GM_E m}{r}$ ?



work

$$\Delta U = - \int_A^B \vec{F} \cdot d\vec{r} = - \int_A^B \frac{GM_E m}{r^2} dr$$

$\uparrow$  Ch. 6       $\uparrow$  Ch. 8  
 Grav. Force

$$= - GM_E m \int_A^B \frac{dr}{r^2} = GM_E m \left( \frac{1}{r} \right)_A^B$$

$$\underbrace{[-\frac{1}{r}]_A^B}_{\text{generic point}}$$

$\Delta U \propto \frac{1}{r} \rightarrow$  zero of potential energy  $\Leftrightarrow B = \infty$   
 (far apart  $\rightarrow$  no attraction)  $\rightarrow$  Ref is at  $\infty$

$$\Delta U = GM_E m \left( \frac{1}{r} \right)_A^\infty = GM_E m \left( \frac{1}{\infty} - \frac{1}{r_A} \right)$$

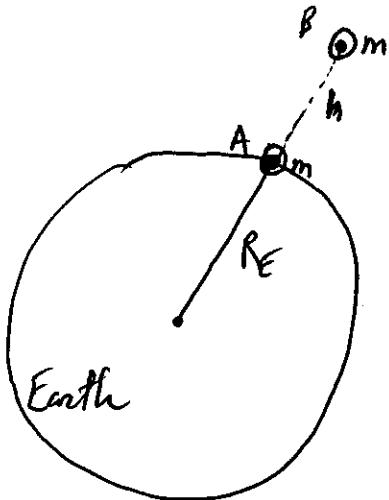
generic point

$$\boxed{\Delta U = - \frac{GM_E m}{r}}$$

higher  $r \rightarrow$  lower potential.

$$\boxed{ME = KE + \Delta U \\ = \frac{1}{2}mv^2 + \frac{GM_E m}{r}}$$

Any connection b/w  $PE = -\frac{GM_E m}{r}$  and  $PE = mgh$ ?



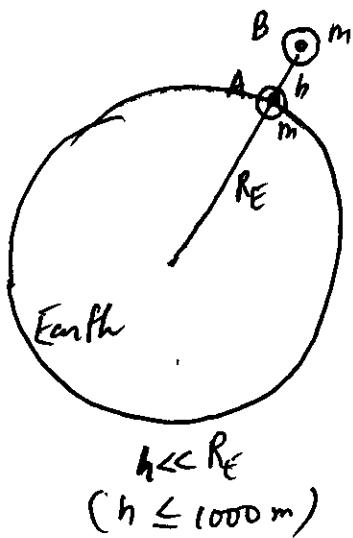
$$\Delta U_{AB} = \Delta PE_{AB} = U_B - U_A$$

$$= -\frac{GM_E m}{R_E + h} + \frac{GM_E m}{R_E}$$

$$= GM_E m \left[ -\frac{1}{R_E + h} + \frac{1}{R_E} \right]$$

$$= GM_E m \left[ \frac{-R_E + R_E + h}{(R_E + h)R_E} \right]$$

$$\Delta U_{AB} = GM_E m \frac{h}{(R_E + h)R_E}$$



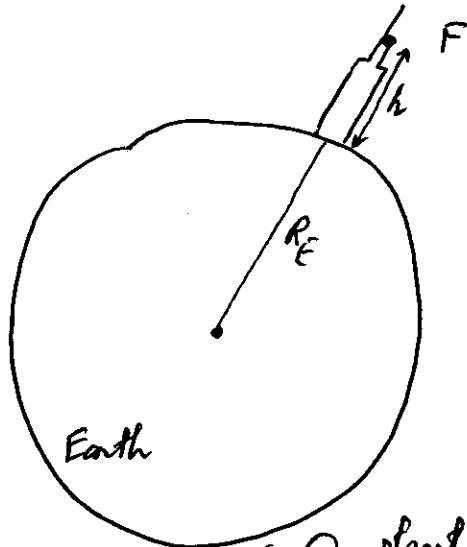
A very good approx:  $R_E + h \approx R_E$

$$\boxed{\Delta U_{AB} = GM_E m \frac{h}{R_E^2} = \underbrace{\frac{GM_E}{R_E^2}}_g mh = mgh}$$

↓  
works for  
smaller heights  
above surface

(8.18)

$$\text{Height of Chicago Sears Tower} : h \longleftrightarrow \Delta g = g - g_h = 0.00136 \frac{\text{m}}{\text{s}^2}$$



$$F = G \frac{M_E m}{r^2} \quad \left\{ \begin{array}{l} @ \text{street level: } g = G \frac{M_E}{R_E^2} \\ @ \text{Top of building: } g_h = G \frac{M_E}{(R_E + h)^2} \end{array} \right.$$

$r$ : center-to-center separation.

$$\begin{aligned} \Delta g = g - g_h &= GM_E \left[ \frac{1}{R_E^2} - \frac{1}{(R_E + h)^2} \right] \\ &= GM_E \left[ \frac{(R_E + h)^2 - R_E^2}{R_E^2 (R_E + h)^2} \right] = GM_E \left[ \frac{2R_E h + h^2}{R_E^2 (R_E + h)^2} \right] \\ &= \underbrace{\frac{GM_E}{R_E^2}}_{\text{no } h} \cdot \underbrace{\frac{2R_E h + h^2}{(R_E + h)^2}}_{\text{with } h} \end{aligned}$$

What is the order of magnitude of  $h = 10\text{m}, 100\text{m}, \sqrt{1000\text{m}}, 1000\text{m}$

$$R_E = 6370000\text{m}$$

$$\text{A very good approximation: } \left\{ \begin{array}{l} R_E + h \approx R_E \\ 2R_E + h \approx 2R_E \end{array} \right.$$

$$\Delta g = \frac{GM_E}{R_E^2} \frac{(2R_E+h)h}{(R_E+h)^2} = \underbrace{\frac{GM_E}{R_E^2}}_{\equiv g} \frac{2R_E \cdot h}{R_E^2}$$

$$\rightarrow h = \frac{\Delta g}{g} \frac{R_E}{2} = \frac{0.00136}{9.81} \cdot \frac{6370000}{2} = 442 \text{ m}$$

8.42

Kepler's Law: "period squared is proportional to the semimajor axis cubed":  $T^2 \propto r^3$

Asteroid Pasachoff:  $\left\{ \begin{array}{l} T_p = 1417 \text{ days} \quad (\text{time to complete one full elliptical orbit around Sun}) \\ r_p ? \quad (r_p \text{ in units of } r_E \Leftrightarrow \frac{r_p}{r_E} ?) \end{array} \right.$

Earth:  $\left\{ \begin{array}{l} T_E \\ r_E \end{array} \right.$

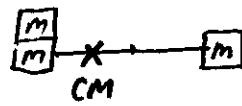
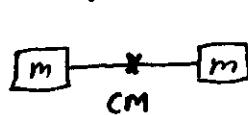
$$\left. \begin{array}{l} T_p^2 \propto r_p^3 \\ T_E^2 \propto r_E^3 \end{array} \right\} \left( \frac{T_p}{T_E} \right)^2 = \left( \frac{r_p}{r_E} \right)^3 \Leftrightarrow \left[ \begin{array}{l} \frac{r_p}{r_E} = \left( \frac{T_p}{T_E} \right)^{2/3} \\ = \left( \frac{1417}{365} \right)^{2/3} \\ = (3.88)^{2/3} = 2.47 \end{array} \right]$$

## Ch 9: System of Particles

So far: we have looked at an object as a point-like object  
 ↳ we took an object as a point located at its center of mass (CM) ↳ Recall the free-body diagrams!

Center of Mass:

average position of all components of a system weighted by their masses.



There are still two positions to average, however the 1st has double weight due to its double mass

Quantitative formulas:

→ Discrete systems

$$\vec{R} = \frac{\sum_i m_i \vec{r}_i}{M}$$

$m_i$ : mass of component  $i$   
 $\vec{r}_i$ : position vector of component  $i$   
 $M$ : total mass of all components  
 $M = \sum_i m_i$

→ Continuous systems:

$$\vec{R} = \frac{\int \vec{r} dm}{M}$$

$dm$ : infinitesimal mass (of the element of mass)  
 $\vec{r}$ : position vector of the element of mass  
 $M$ : total of all elements of mass  
 $M = \int dm$

What is the implication of the components to the 2<sup>nd</sup> Newton's Law of a system (or object)? No, only a subtlety:

- 2nd Newton's Law for a system of particles or components:

$$\vec{F}_{\text{net}} = M \cdot \frac{d^2 \vec{R}}{dt^2}$$

→ No change compared  
to what we have done  
until now

- **Subtlety:** now that we look at components of a system there will be also interactions between these components which are called "internal forces": by 3<sup>rd</sup> Newton's law (action & reaction)



Law (action & reaction)  
The internal force shows up  
in pairs of equal magnitudes  
& opposite directions

$\rightarrow$  sum of all internal forces  
 $\Rightarrow 0$

$\vec{F}_{\text{net}} =$  only involves external forces on the system  
 (no internal forces!)

## Linear Momentum of a system : $\vec{P}$

$$\vec{P} \equiv M \cdot \vec{V} = M \cdot \frac{d\vec{R}}{dt} = M \cdot \frac{d}{dt} \frac{\sum_i m_i \vec{r}_i}{M} = \sum_i m_i \frac{d\vec{r}_i}{dt}$$

Total mass      Velocity of CM      def. of CM  
 of CM    of a discrete system  
 (velocity of component i)

$$\vec{P} = \sum_i \underbrace{m_i \vec{v}_i}_{\vec{p}_i} = \sum_i \vec{p}_i$$

linear momentum  
 of component i

2nd Newton's Law of a system of particles using its total linear momentum :

$$\vec{F}_{\text{net}} = \frac{d\vec{P}}{dt}$$

'Net external force on a system is equal to the change of the total linear momentum over time'

$$\vec{F}_{\text{net}} = 0 \rightarrow \frac{d\vec{P}}{dt} = 0 \text{ or } \vec{P} \text{ is conserved!}$$

Conservation of linear momentum (when there is not a net external force on system)

With this there are 2 conservation laws:

## → Conservation of Mechanical Energy :

Conservation of Mechanical Energy:  
(When only conservative forces are involved, not friction!)

## → Conservation of Linear Momentum

(When there is no net external force on the system).

## Collisions

**Inelastic**: the colliding components stick together after the collision ( $\vec{v}_{1f} = \vec{v}_{2f} = \vec{v}_f$ )

Net external force on system: 0  $\rightarrow$  Conservation of linear momentum:

$$\boxed{m_1 \vec{v}_{1i} + m_2 \vec{v}_{2i} = (m_1 + m_2) \vec{v}_f}$$

before collision      after collision

$ME = KE$  → Kinetic Energy is not conserved:  
 $(10 - 20)$

ME - RE  
(10 or 20)

→ Kinetic Energy is not conserved:

$KE_i - KE_f$  = went into deforming internal structures of either colliding components.

## Elastic

Total KE of colliding components is conserved  
(billiard ball's collisions)

Net external force on system:  $\rightarrow$  Construction of linear Momentum:

$$m_1 \vec{V_{1i}} + m_2 \vec{V_{2i}} = m_1 \vec{V_{1f}} + m_2 \vec{V_{2f}}$$

KE is conserved:

$$\frac{1}{2}m_1v_{1i}^2 + \frac{1}{2}m_2v_{2i}^2 = \frac{1}{2}m_1v_{1f}^2 + \frac{1}{2}m_2v_{2f}^2$$

## More on Elastic Collisions (hard ball collisions)

1D Elastic Collision: we have 2 equations: { conservation of  $\vec{P}$   
conservation of KE

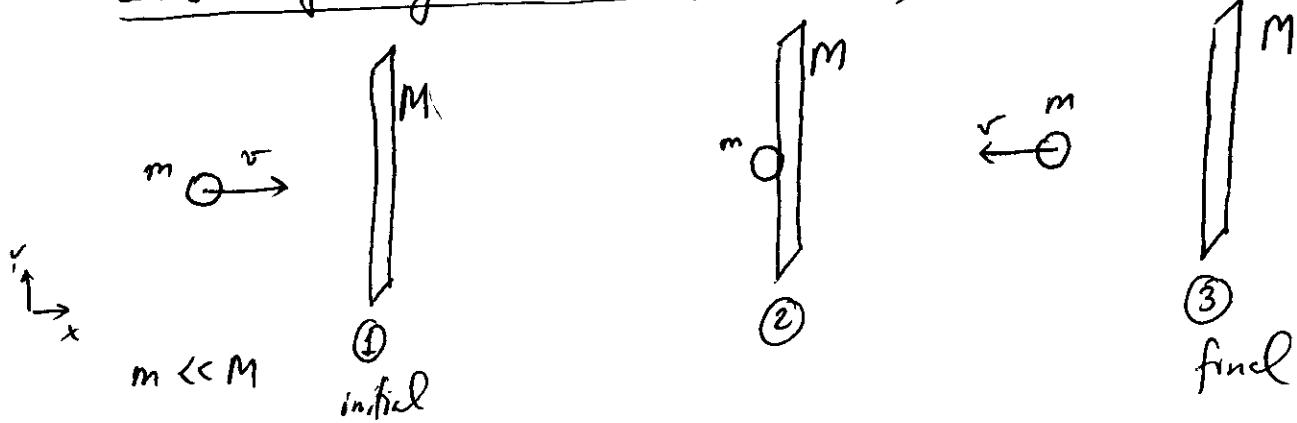
can solve up to 2 unknowns = for example  
given  $m_1, m_2, v_{1i}, v_{2i} \rightarrow$  calculate  $v_{1f}$  &  $v_{2f}$ :

$$\left. \begin{array}{l} 1) v_{1f} = \frac{m_1 - m_2}{m_1 + m_2} v_{1i} + \frac{2m_2}{m_1 + m_2} v_{2i} \\ 2) v_{2f} = \frac{2m_1}{m_1 + m_2} v_{1i} + \frac{m_2 - m_1}{m_1 + m_2} v_{2i} \\ 3) v_{1i} + v_{2f} = v_{1i} + v_{2f} \end{array} \right\} \begin{array}{l} \text{Derived} \\ \text{from} \\ \text{Conservation} \\ \text{of } \vec{P} \text{ & KE} \\ \text{in 1D.} \end{array}$$

2D Elastic Collision: we have 3 equations { conservation of  $\vec{P}_x$   
conservation of  $\vec{P}_y$   
conservation of KE

can solve up to 3 unknowns!  $\rightarrow$  can't solve  
for  $v_{1f}$  &  $v_{2f}$ . A typical problem will  
ask for  $v_{1f}$ ,  $v_{2f}$ ,  $\theta$  (the angle b/w the  
two final velocities)

## Collision of a gas molecule (hard ball) with a container wall



- Wall is much heavier than gas molecule  $\Rightarrow m \ll M$ :

$$\vec{F}_{\text{net, ext}} = 0 \rightarrow \vec{P}_i = \vec{P}_f$$

$$m\vec{v} = m\vec{v}(-\hat{i}) + \vec{\Delta P}_{\text{wall}} \Rightarrow \boxed{\vec{\Delta P}_{\text{wall}} = 2m\vec{v}\hat{i}}$$

After the gas molecule-wall collision, the wall is moving in the  $+\hat{x}$  direction with momentum  $2m\vec{v}$ .

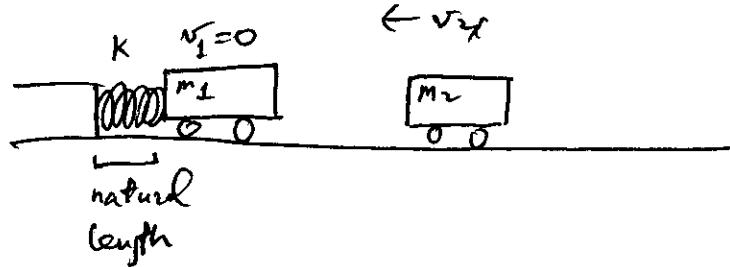
$$\vec{\Delta P}_{\text{wall}} = M \cdot \vec{v}_{\text{wall}} = 2m\vec{v} \hat{i} \rightarrow M v_{\text{wall}} = 2m v$$

$$v_{\text{wall}} = \left(\frac{2m}{M}\right) v$$

Tiny fraction  
(negligible) of  
the molecule  
speed!

### Inelastic Collision

Q.41



initial (before collision)

$$K = 3.2 \times 10^5 \text{ N/m}$$

$$m_1 = 11000 \text{ kg}$$

$$m_2 = 9400 \text{ kg}$$

$$\vec{v}_{2i} = 8.5(-\hat{i}) \frac{\text{m}}{\text{s}}$$



final (after collision)

two cars going together

$$\leftarrow v_f$$

a) C.L.M:  $\vec{P}_i = \vec{P}_f$

$$m_2 \vec{v}_{2i} = (m_1 + m_2) \vec{v}_f$$

$$\vec{v}_f = \frac{m_2}{m_1 + m_2} \vec{v}_{2i} = \frac{9400}{11000 + 9400} 8.5(-\hat{i}) = -3.92 \hat{i} \frac{\text{m}}{\text{s}}$$

Spring will slow the two cars down from  $v_f = -3.92 \hat{i} \frac{\text{m}}{\text{s}}$

to zero :  $KE_{cars} = EPE_{spring}$

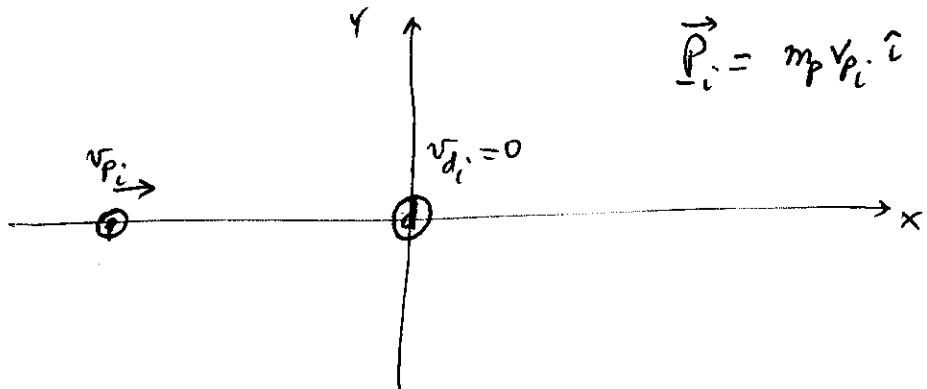
$$\frac{1}{2}(m_1 + m_2) v_f^2 = \frac{1}{2} k (\Delta x)_{max}^2 \rightarrow \begin{cases} \Delta x_{max} = v_f \sqrt{\frac{m_1 + m_2}{k}} \\ = 3.92 \sqrt{\frac{11000 + 9400}{3.2 \times 10^5}} \\ = 0.989 \text{ m} \end{cases}$$

b) Speed of two cars when they rebound from spring:  
gain back some KE from  $EPE_{spring} \rightarrow \vec{v} = 3.92 \hat{i} \frac{\text{m}}{\text{s}}$ .

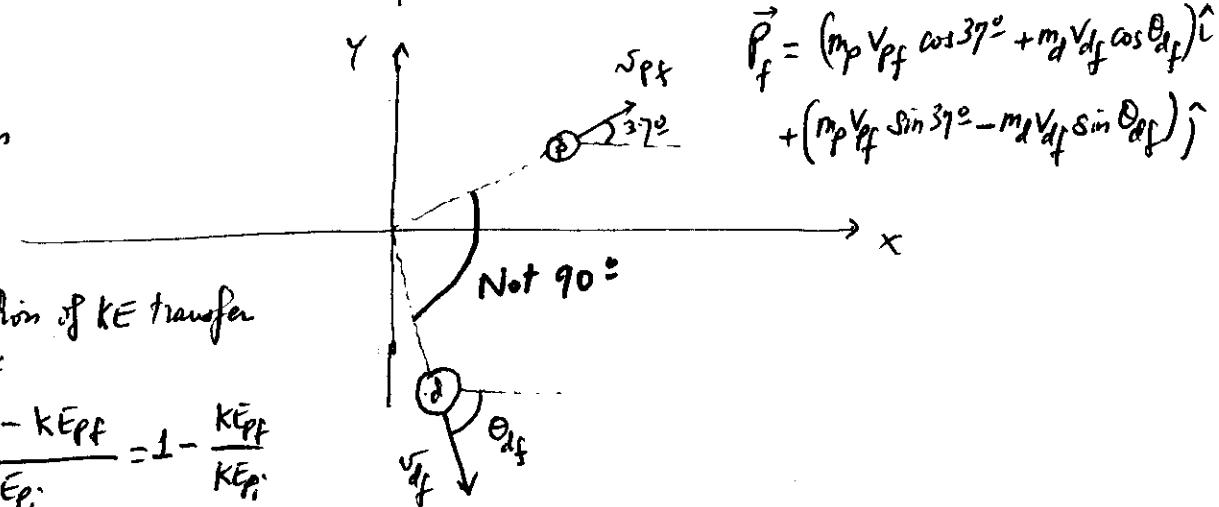
$$\left. \begin{array}{l} \vec{P}_i = \vec{P}_f \\ KE_i = KE_f : KE_{p_i} = KE_{p_f} + KE_{d_f} \rightarrow KE_{d_f} = KE_{p_i} - KE_{p_f} \end{array} \right\}$$

9.65 2D Elastic collision : proton on a stationary deuteron (2u)

Before  
Collision



After  
Collision



Question : fraction of KE transfer from p to d:

$$\frac{KE_{df}}{KE_{p_i}} = \frac{KE_{p_i} - KE_{p_f}}{KE_{p_i}} = 1 - \frac{KE_{p_f}}{KE_{p_i}}$$

$\rightarrow v_{p_f}$  in term of  $v_{p_i}$ :

Note : a) since  $\vec{P}_i$  is in  $+x$  direction, and proton heads  $37^\circ$  (in 1st quad.), the deuteron will head in the 4th quadrant ( $\vec{P}_f$  should have no  $y$ -component)

$$\left. \begin{array}{l} \vec{P}_i = \vec{P}_f \\ m_p v_{p_i} = m_p v_{p_f} \cos 37^\circ + m_d v_{d_f} \cos \theta_{df} \\ 0 = m_p v_{p_f} \sin 37^\circ - m_d v_{d_f} \sin \theta_{df} \end{array} \right\}$$

b)  $m_d = 2 m_p$  ( $37^\circ + \theta_{df} \neq 90^\circ$ , this only happens if  $m_2 = m_1$ )

$$\left. \begin{array}{l} \vec{P}_i = \vec{P}_f \\ v_{p_i} = v_{p_f} \cos 37^\circ + 2v_{d_f} \cos \theta_{df} \quad (1) \\ 0 = v_{p_f} \sin 37^\circ - 2v_{d_f} \sin \theta_{df} \quad (2) \end{array} \right\}$$

Looking for  $v_{p_f}$  in term of  $v_{p_i}$   $\rightarrow$  Need to solve for  $v_{p_f}$ ,  $v_{d_f}$ ,  $\theta_{df}$ : 3 unknowns  $\rightarrow$  write down next  $KE_i = KE_f$ .

$$c) KE_i = KE_f \left\{ \begin{array}{l} \frac{1}{2} m_p v_{pi}^2 = \frac{1}{2} m_d v_{df}^2 + \frac{1}{2} m_p v_{pf}^2 \\ m_d = 2m_p \\ \rightarrow v_{pi}^2 = 2v_{df}^2 + v_{pf}^2 \end{array} \right. \quad (3)$$

From now on use algebraic manipulations to solve for the 3 unknowns:  
 $v_{pf}$ ,  $v_{df}$ ,  $\theta_{df}$  in term of  $v_{pi}$ .

i) Eliminate  $\theta_{df}$ : use  $\cos^2 \theta_{df} + \sin^2 \theta_{df} = 1$

$$\cos \theta_{df} = \frac{(v_{pi} - v_{pf} \cos 37^\circ)}{2v_{df}} \quad (1)$$

$$\sin \theta_{df} = \frac{v_{pf} \sin 37^\circ}{2v_{df}} \quad (2)$$

$$1 = \cos^2 \theta_{df} + \sin^2 \theta_{df} = \frac{1}{4v_{df}^2} \left[ v_{pi}^2 - 2v_{pi}v_{pf} \cos 37^\circ + v_{pf}^2 \underbrace{\cos^2 37^\circ}_{\frac{v_{pf}^2}{v_{df}^2}} + v_{pf}^2 \sin^2 37^\circ \right]$$

$$1 = \frac{1}{4v_{df}^2} \left[ v_{pi}^2 + v_{pf}^2 - 2v_{pi}v_{pf} \cos 37^\circ \right]$$

2) Eliminate  $v_{df}$ : from eq(3)  $v_{df}^2 = \frac{v_{pi}^2 - v_{pf}^2}{2}$

$$2v_{pi}^2 - 2v_{pf}^2 = v_{pi}^2 + v_{pf}^2 - 2v_{pi}v_{pf} \cos 37^\circ$$

$$3v_{pf}^2 - 2v_{pi} \cos 37^\circ v_{pf} - v_{pi}^2 = 0 \quad (ax^2 + bx + c = 0)$$

$$\rightarrow v_{pf} = \frac{2v_{pi} \cos 37^\circ \pm \sqrt{4v_{pi}^2 \cos^2 37^\circ + 12v_{pi}^2}}{6}$$

(Negative sign  
 $v_{pf}$  is negative  
magnitude)

$$v_{pf} = v_{pi} \left[ \frac{2 \cos 37^\circ + \sqrt{4 \cos^2 37^\circ + 12}}{6} \right] = v_{pi} \cdot 0.902$$

Note  $\frac{v_{pf}}{v_{pi}}$  has to be less than one! (some of the proton's energy went into the deuteron after the collision!)

Question: KE transferred from proton to deuteron:

$$\frac{KE_{df}}{KE_{pi}} = 1 - \frac{KE_{pf}}{KE_{pi}} = 1 - \frac{\frac{1}{2} m_p v_{pf}^2}{\frac{1}{2} m_p v_{pi}^2} = 1 - \left( \frac{v_{pf}}{v_{pi}} \right)^2 = 1 - (0.902)^2 = 0.186 \text{ or } 18.6\%$$

For curiosity: we can also solve for  $v_{df}$  &  $\theta_{df}$ :

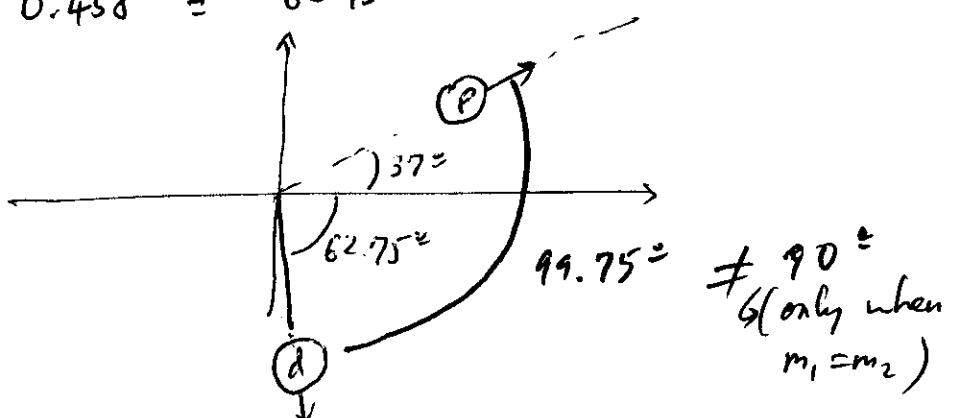
$$3) v_{df}^2 = \frac{v_{pi}^2 - v_{pf}^2}{2} = \frac{v_{pi}^2}{2} \left( 1 - \frac{v_{pf}^2}{v_{pi}^2} \right) = \frac{v_{pi}^2}{2} \underbrace{\left( 1 - 0.902^2 \right)}_{0.186}$$

$$\boxed{v_{df} = v_{pi} \sqrt{\frac{0.186}{2}} = 0.305 v_{pi}}$$

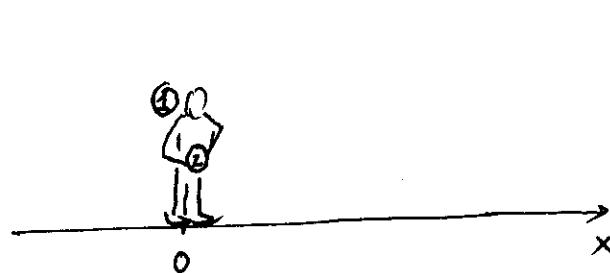
$$1) v_{pi}' = v_{pi} / 0.902 \cos 37^\circ + 2 \cdot 0.305 v_{pi} / \cos \theta_{df}$$

$$\cos \theta_{df} = \frac{1 - 0.902 \cos 37^\circ}{2 \cdot 0.305} = 0.458$$

$$\theta_{df} = \cos^{-1} 0.458 = 62.75^\circ$$

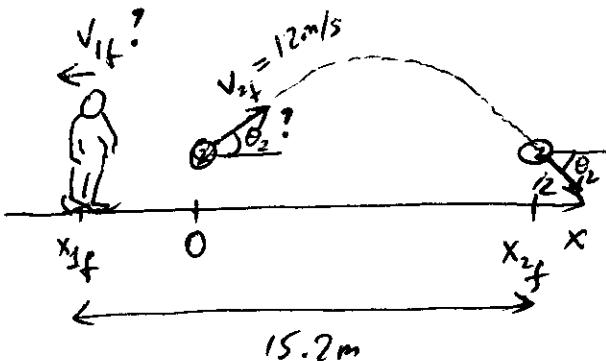


9.52

Initial

$$m_1 = 65 \text{ kg}$$

$$m_2 = 4.5 \text{ kg}$$

Final

Interpretation:  $\vec{P}_i = \vec{P}_f$   
 $0 = \vec{P}_f \rightarrow$  as rock moves in  $+x \rightarrow$  person  
 moves in  $-x$ !

Quantitative analysis:

$$\vec{P}_f = 0 = m_1 \vec{v}_{1f} + m_2 \vec{v}_{2f}$$

(Note: rock followed a 2D not projectile motion!)

Since we are only asked for  $v_{1fx}$ : we'll use the x-component of the conservation of linear momentum equation:

$$0 = m_1 v_{1fx} + m_2 \underbrace{v_{2fx}}_{12 \cos \theta_2} \rightarrow v_{1fx} = -\frac{m_2 \cdot 12 \cos \theta_2}{m_1}$$

$$v_{1fx} = -\frac{4.5 \cdot 12}{65} \cos \theta_2$$

$$\boxed{v_{1fx} = -0.83 \cos \theta_2}$$

Next use  $x_{2f} - x_{1f} = 15.2 \text{ m}$ : via kinematic equations:

Solve for  $x_{2f}$  &  $x_{1f}$  in terms of speeds & angle (or velocity)

Projectile motion:  $x_{2f} = \frac{v_{2f} \cos \theta_2}{v_{2fx}} \cdot 2t_{\text{up}}$  ( $x$ -motion is uniform!)

$t_{\text{up}} = \text{time for rock to go to its max. altitude point:}$

$$v_{2fy} = v_{2f} y_0 - g \cdot t$$

$$0 = v_{2f} \sin \theta_2 - 9.81 \cdot t_{\text{up}}$$

$$t_{\text{up}} = \frac{12 \sin \theta_2}{9.81}$$

$$x_{2f} = 12 \cos \theta_2 \cdot 2 \cdot \frac{12 \sin \theta_2}{9.81}$$

$$= \frac{12^2 \sin 2\theta_2}{9.81}$$

$$(x_{\text{Rup}} = \frac{v_0^2 \sin 2\theta_2}{g})$$

$$x_{2f} = \frac{12^2 \sin 2\theta_2}{9.81}$$

Person's final position:  $x_{1f} = v_{1fx} \cdot 2t_{\text{up}} = -0.83 \cos \theta_2 \cdot 2 \frac{12 \sin \theta_2}{9.81}$   
no friction (on ice)  $\rightarrow$  uniform motion

$$= - \frac{0.83 \cdot (2) \cdot 12}{9.81} \cos \theta_2 \sin \theta_2$$

$$x_{1f} = - \frac{0.83 \cdot 12}{9.81} \sin 2\theta_2$$

person moves in  $-x$  direction!

$$15.2 \text{ m} = x_{2f} - x_{1f} = \frac{12^2 \sin 2\theta_2}{9.81} + \frac{0.83 \cdot 12 \sin 2\theta_2}{9.81} = \frac{12 \sin 2\theta_2}{9.81} (12 + 0.83)$$

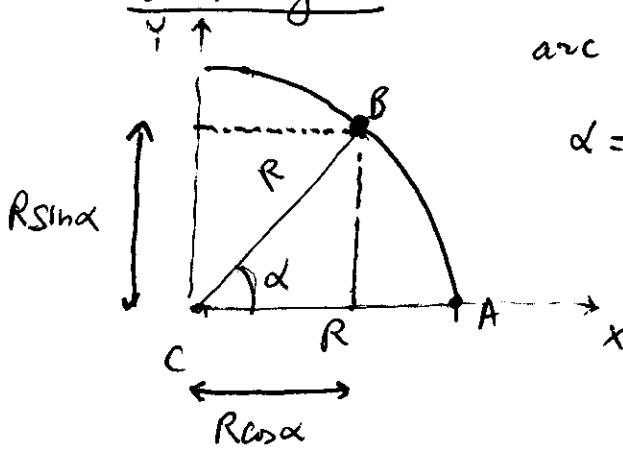
$$\left[ \theta_2 = \frac{1}{2} \sin^{-1} \left( \frac{15.2 \cdot 9.81}{12(12+0.83)} \right) \right] = 37.78^\circ$$

$$V_{fx} = -0.83 \cdot \cos 0_2 = -0.83 \cdot \cos 37.78^\circ = -0.658 \frac{m}{s}$$

↓

result in  $-x$ -direction

About angles:



$$\text{arc } AB$$

$$\alpha = \frac{AB}{R}$$

Position B  $\left\{ \begin{array}{l} (R, \alpha) \rightarrow \text{Polar} \\ \quad \quad \quad \text{coords} \\ (R_{\cos \alpha}, R_{\sin \alpha}) \\ \hookrightarrow \text{Cartesian} \\ \quad \quad \quad \text{coords} \end{array} \right.$

Trigonometric identities:

$$\left\{ \begin{array}{l} \cos^2 \alpha + \sin^2 \alpha = 1 \quad (\text{for any } \alpha) \\ 2 \sin \alpha \cos \alpha = \sin 2\alpha \end{array} \right.$$

(9.57)

' See pg. 108

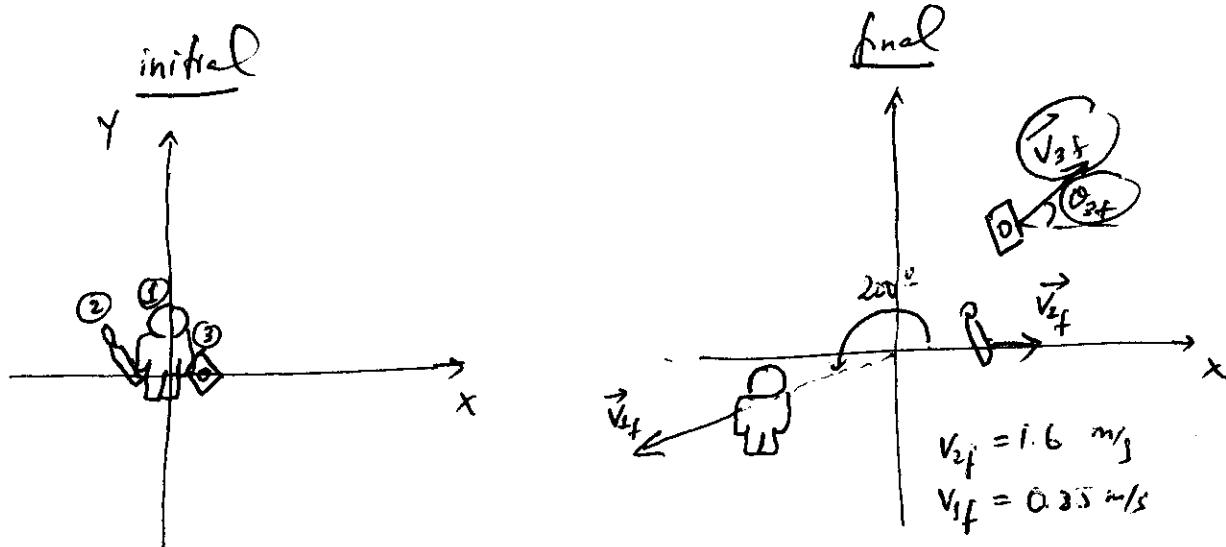
(9.48)

Results at angle  $200^\circ \rightarrow 20$  "million"

Astronaut - oxygen tank - camera

$$m_1 = 60\text{kg} \quad m_2 = 14\text{kg} \quad m_3 = 5.8\text{kg}$$

3 component system



Note: Camera will go off in the 1st quadrant (its x-component along with the O<sub>2</sub>-tank will cancel the x-component of astronaut)

$$\vec{P}_{\text{rel external}} = 0 \rightarrow \vec{P}_i = \vec{P}_f$$

$$0 = m_1 \vec{v}_{1f} + m_2 \vec{v}_{2f} + m_3 \vec{v}_{3f}$$

$$\rightarrow 0 = (m_1 v_{1f} \cos 200^\circ + m_2 v_{2f} + m_3 v_{3f} \cos \theta_{3f}) \hat{i}$$

$$+ (m_1 v_{1f} \sin 200^\circ + m_3 v_{3f} \sin \theta_{3f}) \hat{j}$$

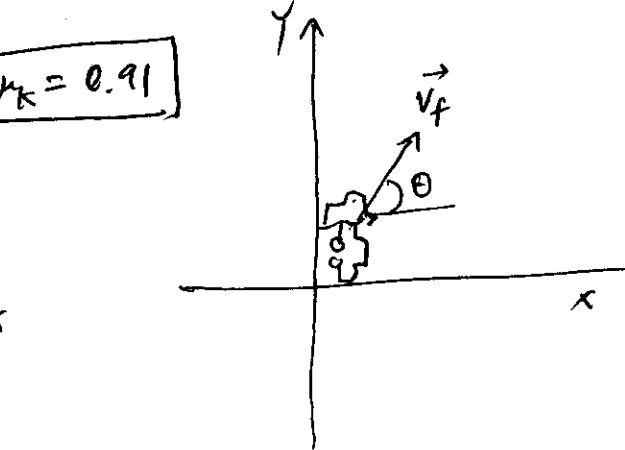
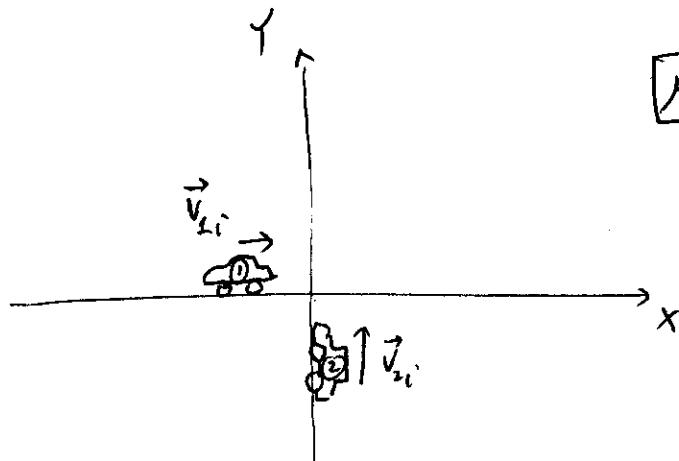
$$\rightarrow \begin{cases} 60 \cdot 0.85 \cdot \cos 200^\circ + 14 \cdot 1.6 + 5.8 \cdot \boxed{v_{3f} \cos \theta_{3f}} \\ 60 \cdot 0.85 \cdot \sin 200^\circ + 5.8 \cdot \boxed{v_{3f} \sin \theta_{3f}} \end{cases} = 0 \rightarrow v_{3fx} = +4.4 \text{ m/s}$$

$$= 0 \rightarrow v_{3fy} = +3 \text{ m/s}$$

2 eqs. & 2 unknowns  $\rightarrow$

$$\vec{v}_{3f} = \sqrt{4.4^2 + 3^2} \hat{m} \rightarrow \begin{cases} v_{3f} = \sqrt{4.4^2 + 3^2} \\ \theta = \tan^{-1} \frac{3}{4.4} \end{cases} \rightarrow \begin{cases} v_{3f} = 5.33 \text{ m/s} \\ \theta_{3f} = 34.3^\circ \text{ (CCW from axis)} \end{cases}$$

9.57

① Toyota  $m_1 = 1200 \text{ kg}$ ② Buick  $m_2 = 2200 \text{ kg}$ 

They lock together after collision  $\rightarrow$  Inelastic Collision  $\rightarrow \vec{v}_{2f} = \vec{v}_{1f} = \vec{v}_f$   
and skid together 22 m.

$$\vec{P}_i = \vec{P}_f$$

$$1) m_1 \vec{v}_{1i} + m_2 \vec{v}_{2i} = (m_1 + m_2) \vec{v}_f$$

2) Skidding will stop when all KE after collision is used up in friction (work done by friction to stop the cars)

$$\frac{1}{2} (m_1 + m_2) v_f^2 = \mu_k (m_1 + m_2) g \cdot x$$

$$1) \begin{cases} m_1 v_{1i} = (m_1 + m_2) v_f \cos \theta \\ m_2 v_{2i} = (m_1 + m_2) v_f \sin \theta \end{cases} \rightarrow 1200 v_{1i} = 3400 v_f \cos \theta \\ 2200 v_{2i} = 3400 v_f \sin \theta$$

$$2) \frac{1}{2} 3400 v_f^2 = 0.91 (3400) \cdot 9.81 \cdot 22$$

$$1200^2 v_{1i}^2 + 2200^2 v_{2i}^2 = 3400^2 v_f^2 \quad (\cos^2 \theta + \sin^2 \theta = 1)$$

$$1200^2 v_{1i}^2 + 2200^2 v_{2i}^2 = \underbrace{3400^2 \cdot 2 \cdot 0.91 \cdot 9.81 \cdot 22}_{1.44 \times 10^6 \quad 4.84 \times 10^6}$$

$$v_{1i}^2 + 3.36 v_{2i}^2 = 3.15 \times 10^3 \rightarrow \frac{4.54 \times 10^9}{25 \text{ km}} \text{ is } 6.94 \text{ m} \quad \text{Clearly this equation is not satisfied if both } v_{1i} \text{ and } v_{2i} \text{ are } 6.94 \text{ or less!}$$

9.67

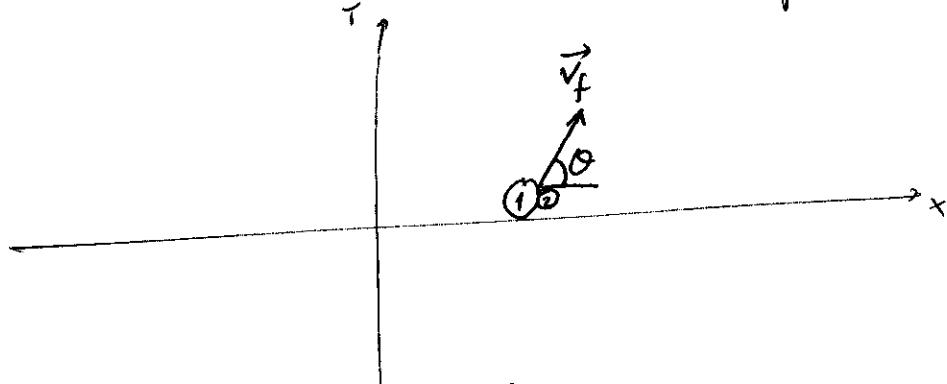
$$m_1 = 32 \text{ u}$$

$$m_2 = 16 \text{ u}$$

$$\vec{v}_{1i} = 580 \text{ m/s} \hat{i}$$

$$\begin{aligned}\vec{v}_{2i} &= 870 \frac{\text{m}}{\text{s}} @ 27^\circ \\ &= 870 \cos 27^\circ \hat{i} + 870 \sin 27^\circ \hat{j}\end{aligned}$$

final

Inelastic collision:  $\vec{P}_i = \vec{P}_f$ 

$$m_1 \vec{v}_{1i} + m_2 \vec{v}_{2i} = (m_1 + m_2) \vec{v}_f$$

$$m_1 = 2m_2$$

$$2 \cdot 580 \hat{i} + 870 \cos 27^\circ \hat{i} + 870 \sin 27^\circ \hat{j} = 3(v_{fx} \hat{i} + v_{fy} \hat{j})$$

$$\begin{aligned}2 \cdot 580 + 870 \cos 27^\circ &= 3v_{fx} \rightarrow v_{fx} = 645.06 \text{ m/s} \\ 870 \sin 27^\circ &= 3v_{fy} \rightarrow v_{fy} = 131.66 \text{ m/s}\end{aligned}$$

$$\rightarrow v_f = \sqrt{645^2 + 131.66^2} = 658.4 \text{ m/s}$$

$$\theta = \tan^{-1} \frac{131.66}{645.06} = 11.54^\circ$$

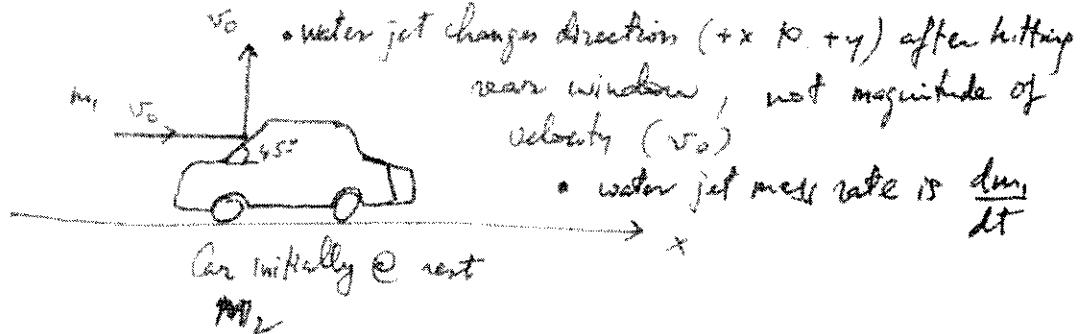
104

$$\frac{1}{2} m v_{\min}^2 = m g 17 \left( 1 - \cos \left( \sin^{-1} \frac{10}{17} \right) \right)$$

$$v_{\min} = \sqrt{2 \times 9.81 \times 17 \left( 1 - \cos 36^\circ \right)} = 7.98 \text{ m/s}$$

(9.43)

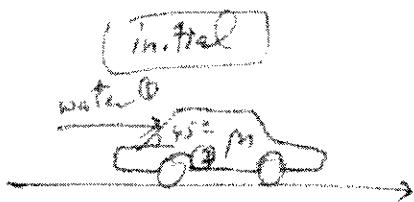
• No friction

a)  $a_x$  for the car?

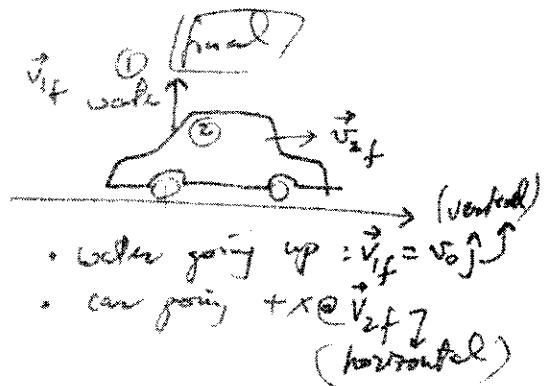
→ The collision b/w jet of water & car @ rest transfers some of its momentum into the car allowing it to go from zero speed to non-zero speed : requiring an acceleration in the horizontal direction  $\rightarrow a_x$ .

→  $\vec{F}_{\text{ext}} = 0$  all forces are b/w components (water jet & car) of the system. (No friction).

$$\hookrightarrow \vec{P}_i = \vec{P}_f$$



- water jet  $\frac{dm}{dt}$  @
- $\vec{v}_i = v_0 \hat{i}$
- car @ rest, mass  $M_2$



$$m_1 \vec{v}_{1i} = m_1 \vec{v}_{1f} + m_2 \vec{v}_{2f}$$

Note: water & car no. t sticking together after collision!

1195

$$\text{observation: } \vec{a} = \frac{d\vec{v}}{dt} \rightarrow \text{or } \vec{a}_x = \frac{d\vec{v}_x}{dt}$$

$$\rightarrow \vec{v}_{if} = m_1 \frac{1}{m_2} (\vec{v}_{ic} - \vec{v}_{if}) = m_1 \frac{1}{m_2} (v_0 \hat{i} - v_0 \hat{j})$$

$$\rightarrow \vec{a} = \frac{d\vec{v}_x}{dt} = \frac{dm_1}{dt} \frac{v_0}{m_2} (\hat{i} - \hat{j})$$

↓  
only  $m_1$   
is changing

Observation: final acceleration of car has 2 components.

$$\left\{ \begin{array}{l} a_x = \frac{dm_1}{dt} \frac{v_0}{m_2} \\ \end{array} \right. \quad \left. \begin{array}{l} \text{forward in +x.} \end{array} \right.$$

$$a_y = - \frac{dm_1}{dt} \frac{v_0}{m_2} \quad \begin{array}{l} \text{(since initially there} \\ \text{was no momentum} \\ \text{in the y direction,} \\ \text{and since finally} \\ \text{the water goes up} \\ \rightarrow \text{car gets pushed} \\ \text{down)} \end{array}$$

b) Max. speed reached by the car?

Why max? or can we accelerate the car to  
a speed with this water jet? No b/c when  
the car reaches  $v_0$  (same speed as the water jet)  
no more pushing or momentum transfer is possible.

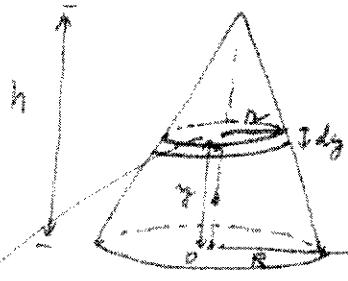
→ max speed for car is  $v_0$ !

(172)

9.39

Y

x



infinitesimal disk of thickness  $dy$  of mass  $dm$ ,  
radius  $r$ , located @ distance  $y$  above  
the base.

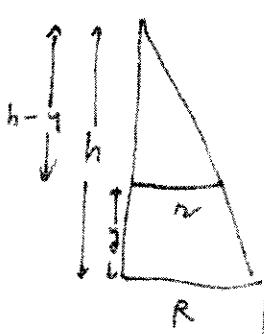
$$\vec{R} = \frac{1}{M} \int \vec{r} dm \quad \begin{cases} x_{cm} = \frac{1}{m} \int x dm \\ y_{cm} = \frac{1}{m} \int y dm \end{cases}$$

total mass.

position vector of center of mass.

b/c symmetry  $x_{cm} = 0$  in the  
coord. syst. shown. ( $Y$  axis  
coincide w/ the axis of  
symmetry of the cone).

$$\rightarrow y_{cm} = \frac{1}{M} \int y dm = \frac{1}{M} \int_0^h y \rho \pi R^2 \left(1 - \frac{y}{h}\right)^2 dy =$$



$$\frac{r}{h-y} = \frac{R}{h} \rightarrow r = R \frac{(h-y)}{h}$$

$$r = R \left(1 - \frac{y}{h}\right)$$

Infinitesimal disk volume is  
 $dV = \pi r^2 dy$

$$\text{since } \rho = \frac{dm}{dV} \rightarrow dm = \rho dV$$

$$dm = \rho \pi r^2 dy = \rho \pi R^2 \left(1 - \frac{y}{h}\right)^2 dy$$

$$\downarrow y_{cm} = \frac{\rho \pi R^2}{M} \int_0^h y \left(1 - \frac{2y}{h} + \frac{y^2}{h^2}\right) dy$$

$$= \frac{\rho \pi R^2}{M} \int_0^h \left(y - \frac{2}{h} y^2 + \frac{1}{h^2} y^3\right) dy$$

$$\int x^n dx = \frac{x^{n+1}}{n+1}$$

$$\boxed{\rho = \frac{M}{\text{Vol of cone}}} \quad \downarrow \quad = \frac{3}{h} \left[ \frac{y^2}{2} - \frac{2}{h} \frac{y^3}{3} + \frac{1}{h^2} \frac{y^4}{4} \right]_0^h$$

$$\boxed{\rho = \frac{M}{\pi R^2 h / 3}} \quad \downarrow \quad = \frac{3}{h} \left[ \frac{h^2}{2} - \frac{2}{3} h^2 + \frac{1}{4} h^2 \right] = 3h \left[ \frac{1}{2} - \frac{2}{3} + \frac{1}{4} \right]$$

$$= 3h \left[ \frac{6-8+3}{12} \right] = \frac{3h}{12} \cdot \frac{h}{4}$$